

Intro, (93) Well done, even if not concise.

(1)

Pedlosky's (1987) equation (6.8.9)

$$\frac{d_0}{dt} \left[ \xi_0 + \beta y - \frac{\partial}{\partial z} \frac{p_0}{s} \right] = 0$$

derives from the <sup>...precisely</sup> ~~exact~~  $O(1)$  vorticity equation ... assuming a flat earth with the Coriolis parameter varying linearly with latitude (6.3.17)

$$\frac{d_0}{dt} \left( \xi_0 + \beta y \right) = \frac{1}{\rho_s} \frac{\partial}{\partial z} (\rho w_0)$$

The latter equation represents quasi-geostrophic dynamics in its full representation, the former equation is an approximation of the latter. Both equations, their validity, and physical meaning and interpretation are discussed subsequently.

The independent variables are  $(x, y, z, t)$ , the dependent variables

are  $\rightarrow \xi_0 = \frac{\partial u_0}{\partial x} - \frac{\partial v_0}{\partial y}$  i.e.  $v_0, u_0$  horizontal geostrophic velocities

$\rho = \rho_s(z) + \rho'(x, y, z, t)$  and  $\rho' = \rho_0 + \epsilon \rho_1 + \dots$

$\rho_s$  is a known background density state and  $\rho'$  is a density perturbation due to the motion.

I doubt

The horizontal velocities to be geostrophic means that to  $O(1)$  in an  $\epsilon$ -expansion ( $\epsilon = U/L \ll 1$ , the Rossby #)

$(u_0, v_0, p_0)$  satisfy

$$v_0 = \frac{\partial p_0}{\partial x}, \quad u_0 = -\frac{\partial p_0}{\partial y} \quad \checkmark$$

and the vertical hydrostatic equation

$$p_0 = -\frac{1}{\rho_s} \frac{\partial}{\partial z} (\rho_s p_0) \quad \checkmark$$

Pedlosky (6.3.6)

Hence  $\mathcal{O}_0 = \nabla_H p_0$ , i.e. all three dependent variables can be expressed in terms of  $p_0$ , the first order pressure perturbation.

The expression for thermodynamic processes enters the dynamical description when the  $O(\epsilon)$  vertical velocity  $w_1$  in (6.3.17) is expressed in terms of  $p_0$  only.  $\gamma$  and  $\beta$  are nondimensional parameters discussed later.

Then, (6.8-9) is a nonlinear (because of  $\frac{d\mathcal{O}}{dt} = \frac{\partial}{\partial t} + u_0 \frac{\partial}{\partial x} + v_0 \frac{\partial}{\partial y}$ )

partial differential equation in  $(x, y, z, t)$  for the depend variable  $p_0$  only. All other dependent quantities can then be solved for once  $p_0$  is known.

This question has a one line answer: Pot. vort. (or angular momentum) is conserved following the hor. motion.

(1)

The equation

$$\frac{d_0}{dt} \left[ \xi_0 + \beta y - \frac{\partial}{\partial z} \frac{p_0}{S} \right] = 0$$

is a conservation statement following a geostrophically ~~stream~~ <sup>conservation of what?</sup> moving particle, i.e. for a particle moving with  $(u_0, v_0)$  the quantity

$$\xi_0 + \beta y - \frac{\partial}{\partial z} \frac{p_0}{S} \quad \text{is} \quad \frac{\Sigma}{\delta}$$

is a constant and stays constant for that particle as it moves.

The first term  $\xi_0$  is its relative vorticity, i.e. its ~~total vorticity~~  $\left( \xi_0 = \vec{k} \cdot \nabla \times \vec{u}_0 \right) = \left( \frac{\partial v_0}{\partial x} - \frac{\partial u_0}{\partial y} \right)$  local vorticity.

The second term  $\beta y$  represents the planetary vorticity, i.e. the effects of the earth's rotation upon the particle, and the last term  $-\frac{\partial}{\partial z} p_0/S$  represents vortex line stretching or squashing due to variations of isopycnal surfaces.

Hence it is the last term which introduces the effects of a variable density stratification.

Because the equation is ~~too~~ scaled it states (for ~~the~~  $\beta = O(1)$  and  $S = O(1)$ ) that the contribution of ~~vorticity~~ a particle's vorticity, by relative vorticity, by absolute vorticity, and by vortex line stretching due to density variations are of the same importance and its sum stays constant as the particle moves with  $(u_0, v_0)$

(2) The assumptions to derive the precisely quasi-geostrophic potential vorticity equation (6.3.17) are:

(i) Small Rossby #  $\epsilon \ll 1$

$$\epsilon = \frac{U}{2\Omega L} \ll 1 \quad \text{meaning?}$$

(ii)  $\beta$ -plane, i.e. planetary vorticity varies linearly with latitude

$$\frac{L}{r_0} = \mathcal{O}(\epsilon)$$

$r_0$  the radius of the earth  
 $L$  a lengthscale of the motion

(iii) Vortex tube stretching due to the free surface is negligible

$$F = \left(\frac{L}{R_0}\right)^2 = \left(\frac{L^2}{\sqrt{gD} |f_0|}\right)^2 = \mathcal{O}(\epsilon)$$

$\sqrt{gD} |f_0|$  is the <sup>barotropic</sup> Rossby radius of deformation,  $D$  is a depth scale,  $f_0 = 2\Omega$  the local Coriolis part of earth's rotation

(iv) We stay away from frictional boundaries, i.e. interior dynamics only:

$$(E_V, E_H) = \mathcal{O}(\epsilon)$$

$E_V, E_H$  are vertical and horizontal Ekman numbers

(v) To ensure hydrostatic balance, we require  $\delta = \frac{D}{L} \ll 1$

Another additional assumption is needed to derive equation (3.8.9) from the precise continuity equation (3.3.17):

(vi) The scale of motion is of the order of the baroclinic Rossby radius of deformation  $L_D$  (mesoscale, synoptic scale), i.e.,

$$S = \left( \frac{L_D}{L} \right)^2 = \left( \frac{|g' D^*|}{f_0 L} \right)^2 = O(1)$$

$g'$  is a reduced gravity  $\frac{\Delta \rho}{\rho} g$  indicating the gravity effects of different density 'layers', say,  $D$ , then, is a typical layer depth rather than the total depth.

(vii) Neglect internal heating ✓

Crucial in facilitating the thermodynamics to find a simple model of how the thermodynamics enter the quasi-geostrophic dynamics is condition (vi), which can also be restated that the buoyancy frequency  $N_s$  (Brunt-Väisälä frequency) is small

$$N_s = \left( -\frac{g}{\rho} \frac{\partial \rho_s}{\partial z_s} \right)^{1/2} \quad \text{in dimensional form}$$

$$\text{also, } D/H_* \ll 1$$

$$\text{and } h_B/D \leq O(\epsilon)$$

$$\text{as } S^{1/2} = \frac{N_s D}{f_0 L} \quad \checkmark$$

however, the dynamical important length scale ( $S=O(1)$ ) is  $L_D$ , the mesoscale.

28  
30

(3) The physical interpretation of  $\frac{d\sigma}{dt} \frac{\sigma_0}{\sigma_0}$  is

the geostrophic rate of change of the relative vorticity  $\frac{\sigma_0}{\sigma_0}$ , i.e., to be important the horizontal geostrophic velocities have to change rapidly, <sup>enough</sup> because potentially this term might be balanced by the  $\beta$  effect, i.e.  $\beta \frac{d\sigma}{dt} y$  that is as the individual particle moving with  $(u_0, v_0)$  changes its latitude (its position northward,  $y$ ) its planetary vorticity increases, <sup>or</sup> decreases. On the oceanic scale,  $L \sim 2000$  km we usually neglect the <sup>relative</sup> relative vorticity, to be small, but here ~~we~~ our spatial length scale is ~~different~~ (baroclinic Rossby radius, instead of the barotropic one) and therefore the relative vorticity is an  $O(1)$  contributor in the vorticity balance (potentially). The dynamical flow field resembling rapid spatial changes is an eddy field, the particles may enter and leave eddies, changing their <sup>rel.</sup> vorticity according to their position. The term  $-\frac{1}{S} \frac{d\sigma}{dt} \left( \frac{\partial p_0}{\partial z} \right)$  indicates that relative and planetary

vorticity might be balanced by changing ~~the~~ isopycnals ( $p_0 = \text{const}$ ) vertically as the fluid particle moves along. It experiences vorticity like stretching and squashing because ( $p_0 = \text{const}$ ) is a ~~three dimensional~~ surface varying spatially its depth. The particles with a certain density stays always on its isopycnal surface, hence it has to follow the up and

downs of that surface. Here it comes out quite clearly that  
the variable isopycnal induces vertical motion  $w$ , (see eqn. 6.3.17)  
i.e. water table stretching. ✓

A good, but not concise, account.

12  
—  
12

(4) As it was mentioned in the introduction the three dependent variables  $(u_0, v_0, p_0)$  can be expressed in terms of the pressure perturbation  $p_0$  alone. Once  $p_0$  is found as  $p_0(x, y, z, t)$  the geostrophic equations for a stratified fluid can be used to find the quantities  $(u_0, v_0)$  and the hydrostatic equation is used to find  $p_0$ . These set of diagnostic equations is solved trivially once the vorticity equation is solved.

Density variations eventually remove the two-dimensional character of barotropic quasi-geostrophic flow inducing vertical variability.

This is expressed in the geostrophic equations which can be rewritten in terms of vertical shears utilising the hydrostatic equation. The resulting set of equations are the thermal wind or (Gill; 1982) relative geostrophic currents. They are diagnostic relationships of vertical geostrophic current shears and horizontally varying density field:

$$\frac{\partial u_0}{\partial z} = \frac{\partial v_0}{\partial y}, \quad \frac{\partial v_0}{\partial z} = -\frac{\partial p_0}{\partial x}$$

Pedlosky (6.8.7.)

or dimensionally

$$f \frac{\partial u}{\partial z} = \frac{g}{\rho} \left( \frac{\partial \rho}{\partial y} \right)_p, \quad f \frac{\partial v}{\partial z} = -\frac{g}{\rho} \left( \frac{\partial \rho}{\partial x} \right)_p \quad \text{Gill (2.7.9)}$$

"...where the derivatives on the right hand side are taken on constant-pressure surface..." Gill (1982, p. 214), i.e. a particle moving on its isopycnal. All velocities are first order geostrophic velocities



(4ff)

The continuity equation to this order of approximation  $O(1)$  is still degenerated as

$$\frac{1}{\rho_s} \frac{\partial}{\partial z} (\rho_s w_0) + \frac{\partial u_0}{\partial x} + \frac{\partial v_0}{\partial y} = 0$$

Redlosky  
(6.3.7)

but the geostrophic velocities are nondivergent:

$$\frac{\partial u_0}{\partial x} + \frac{\partial v_0}{\partial y} = 0 \quad \checkmark$$

Therefore

$$\frac{\partial}{\partial z} \rho_s w_0 = 0 \quad \checkmark \quad \text{also}$$

and because  $w_0$  vanishes at the top and the bottom  $w_0 = 0 = w_0(z)$  we still have a stream function which is — in essence — the pressure perturbation  $p_0$ .

Another diagnostic equation for known  $p_0, u_0, v_0, \rho_0$  can give us the  $O(\epsilon)$  vertical velocity field  $w_1$  in the absence of thermal heating:

$$-\frac{d_0 \rho_0}{dt} + w_1 S = 0 \quad \checkmark \quad (6.8.3)$$

All quantities are known except  $w_1$ .

It is noted in passing that the thermal wind equations show that vortex tube tilting might be important too. The vertical ~~shear~~ shear of the (geostrophic) horizontal velocities induces vorticity not in the vertical but in the horizontal ~~plane~~. But it doesn't enter. However, the thermal wind equations state that horizontal ~~to~~  $O(1)$  vort. dynamics. density variations induce vertical current shear and possibly current vector rotation ~~also~~ with depth.

22  
22

(5) A particle of fluid approaching the Gulf Stream from the subtropical gyre has a certain amount of potential vorticity which it obtained from the wind stress curl. The major (or gyre scale) vorticity balance in the interior is the Sverdrup balance, i.e. wind stress curl is balanced by planetary vorticity in a barotropic ocean. The length scales for such processes are the barotropic Rossby radius. The dynamical length scale for which the dynamics represented by eqn. (3.8.9) hold is the baroclinic Rossby radius  $L_b$  which is usually much smaller than the barotropic radius (e.g.  $\ll$  in the ocean). Therefore, I think I can assume isopycnals which ~~are~~ do not vary along latitudes to enter the Gulf Stream region. The length scale of variation of the two regions is very different, so (3.8.9) variations on the large (gyre) scale are not seen or negligible on the small (mass, or eddy) scale.

$L$  is the scale of the WSC

OK

OK

Let's consider for a moment the steady version of  $\frac{d\sigma}{dt}$  which

is  $u_0 \frac{\partial}{\partial x} + v_0 \frac{\partial}{\partial y}$  to simplify the argument. In the high velocity regions of the Gulf Stream the relative vorticity (which was

small for the particles coming from the subtropical gyre) is large when compared to the planetary vorticity. The only term which can possibly ~~support~~ <sup>balance</sup> (apart of friction) the relative vorticity there is vorticity stretching due to changes of the upper layer depth

not in eqn (3.8.9)  
 same as eqn (6.8.9)

or strongly sloping isopycnals. The particles approaching the Gulf Stream ~~but~~ gain anticyclonic relative vorticity which is negative vorticity.



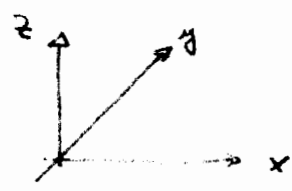
To balance this negative vorticity tendency the stretching term

$-\frac{\partial}{\partial z} \rho_0$  has to provide the necessary positive vorticity tendency

hint is for the moving fluid column  $\rightarrow \frac{\partial \rho_0}{\partial z} > 0$  &  $\frac{\partial \rho_0}{\partial x} < 0$

$\frac{\partial \rho_0}{\partial z}$  has to ~~decrease~~ increase in magnitude as in a stably stratified fluid  $\frac{\partial \rho}{\partial z} < 0$

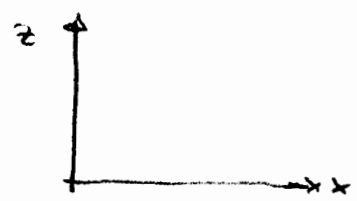
Because of  $\frac{\partial v_0}{\partial z} = -\frac{\partial u_0}{\partial x}$



is larger than zero:

$\frac{\partial v_0}{\partial z} > 0$

$\frac{\partial \rho_0}{\partial x} < 0$ , hence the density increases from the east toward the west



The relative vorticity which is small for the fluid column entering

the gulf stream ~~in some~~ increases and enters the vorticity balance.

This <sup>anti-cycl.</sup> relative vorticity is balanced by vortex tube <sup>squeezing</sup> stretching due to the rising isopycnals. Vertical velocities, therefore, are relatively large  $O(\epsilon)$ . Another interpretation is that available potential energy from the subtropical gyre is released in form of kinetic energy because of the rise of the isopycnal surfaces. ✓ good!

Adding the time dependence the qualitative result of waves, instabilities and eddies due to this transfer of vorticity is anticipated. This, also, is indicated by the scale for which these dynamics hold: the baroclinic Rossby radius of deformation which corresponds closely to the eddy scale in the ocean.

In brief: In a stratified ocean which over most of its interior is in Sverdrup balance the ocean side of the Gulf Stream can be explained qualitatively by a vorticity transfer from the density field to the flow field, i.e. vortex tube squashing due to rising isopycnals can (partly) provide ~~vorticity~~ relative vorticity present in the Gulf Stream. ✓

Very well done!

13  
—  
12

(6) A simplified <sup>first</sup> version of

$$\frac{d_0}{dt} \left[ \xi_0 + \beta y - \frac{1}{S} \frac{\partial p_0}{\partial z} \right] = 0$$

which is appropriate to the interior of the subtropical gyre is certainly the steady part of it

$$u_0 \frac{\partial}{\partial x} \left( \xi_0 + \beta y - \frac{1}{S} \frac{\partial p_0}{\partial z} \right) + v_0 \frac{\partial}{\partial y} \left( \xi_0 + \beta y - \frac{1}{S} \frac{\partial p_0}{\partial z} \right) = 0$$

? why steady?

or

$$u_0 \frac{\partial}{\partial x} \left( \xi_0 - \frac{1}{S} \frac{\partial p_0}{\partial z} \right) + v_0 \beta + v_0 \frac{\partial}{\partial y} \left( \xi_0 - \frac{1}{S} \frac{\partial p_0}{\partial z} \right) = 0$$

The relative vorticity is small compared to the planetary vorticity, <sup>and also on the mesoscale.</sup> ~~and~~ in the interior of the gyre leaving

$$\frac{u_0}{S} \frac{\partial^2}{\partial z \partial x} p_0 + \frac{v_0}{S} \frac{\partial}{\partial z} p_0 = v_0 \beta$$

$$u_0 \frac{\partial^2}{\partial z \partial x} p_0 + v_0 \frac{\partial^2}{\partial z \partial x} p_0 = v_0 \frac{\beta}{S}$$

The correct statement is that the gradient of  $S_0$  is small compared to that of the plan. vort.,  $\nabla S_0 \ll \beta$  or  $\beta \gg 1$ .

$$\frac{d_0}{dt} \frac{\partial p_0}{\partial z} = v_0 \frac{\beta}{S} \quad \text{or} \quad \frac{d_0}{dt} \left( \beta y - \frac{1}{S} \frac{\partial p_0}{\partial z} \right) = 0$$

I have never seen such a balance, because the vortex tube stretching in the ocean's interior is usually provided by the Ekman pumping from an Ekman layer forced by the wind stress <sup>This is not the interior vorticity</sup> ~~work~~. Here the planetary vorticity is balanced by vortex tube stretching, etc, and due to sloping isopycnals. The vorticity <sup>The net effect is still governed by WSC.</sup> "pumping" velocity  $w_e = O(\epsilon)$  would drive such a stratified, unforced ocean. Well, unforced is not quite the right word as respect

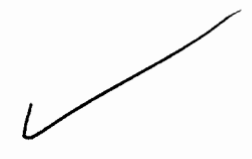
OK a possible forcing from the boundary now is differential heating and cooling at the surface. <sup>or WSC</sup> And indeed this is the case as the tropics receive more sunlight, heat than the mid latitudes. This causes sloping North-South isopycnals. Subduction of denser water and provides of fluid columns on them could eventually be described by this simplified ~~version~~ (i.e. steady) version of (6.8.8).

OR Note that the boundary pumping time scales with  $S^{-1}$  where  $S = \left(\frac{L_0}{L}\right)^2$ , but on the gyre scale  $L_0 \ll L$ , hence ~~S~~ rather  $S \ll 1$ .  $S$  is no longer  $O(1)$ . Thus rather than (6.8.9) <sup>equation (6.8.4)</sup> is to be used.   
 no, your eq. is the one.

Another interpretation of  $S^{-1}$  is that it is the ratio between available potential energy to kinetic energy, i.e.,  $S = O(1)$  gives the length scale  $L$  on which ~~all the~~ ~~the~~ APE associated with the sloping isopycnals is <sup>equally partitioned with</sup> ~~transformed into~~ kinetic energy. In the subtropical gyre we found that  $S = \left(\frac{L_0}{L}\right)^2 \ll 1$ .   
 ~~these results~~ that only a very small amount of

I conclude that only a very small part of the total energy is kinetic

$$S''_c = \frac{L_c}{L} = \frac{\text{kinetic energy}}{\text{HPE}}$$



The HPE energy is interpreted as the part of the potential energy (stored in the density, mass distribution) which can be transformed into kinetic energy supporting motion due to sloping isopycnals. As Gill (1982, p. 200) pointed out "... potential energy is hard to extract ~~from~~ in a rotating frame."

However, the ~~too~~ vorticity balance given in this section and its energy considerations (kinetic energy context) probably resembles some features of the thermohaline circulation & speculate. also the wind driven part. (See Ped., Chap. 6, on gyre scale circ.)

13  
16